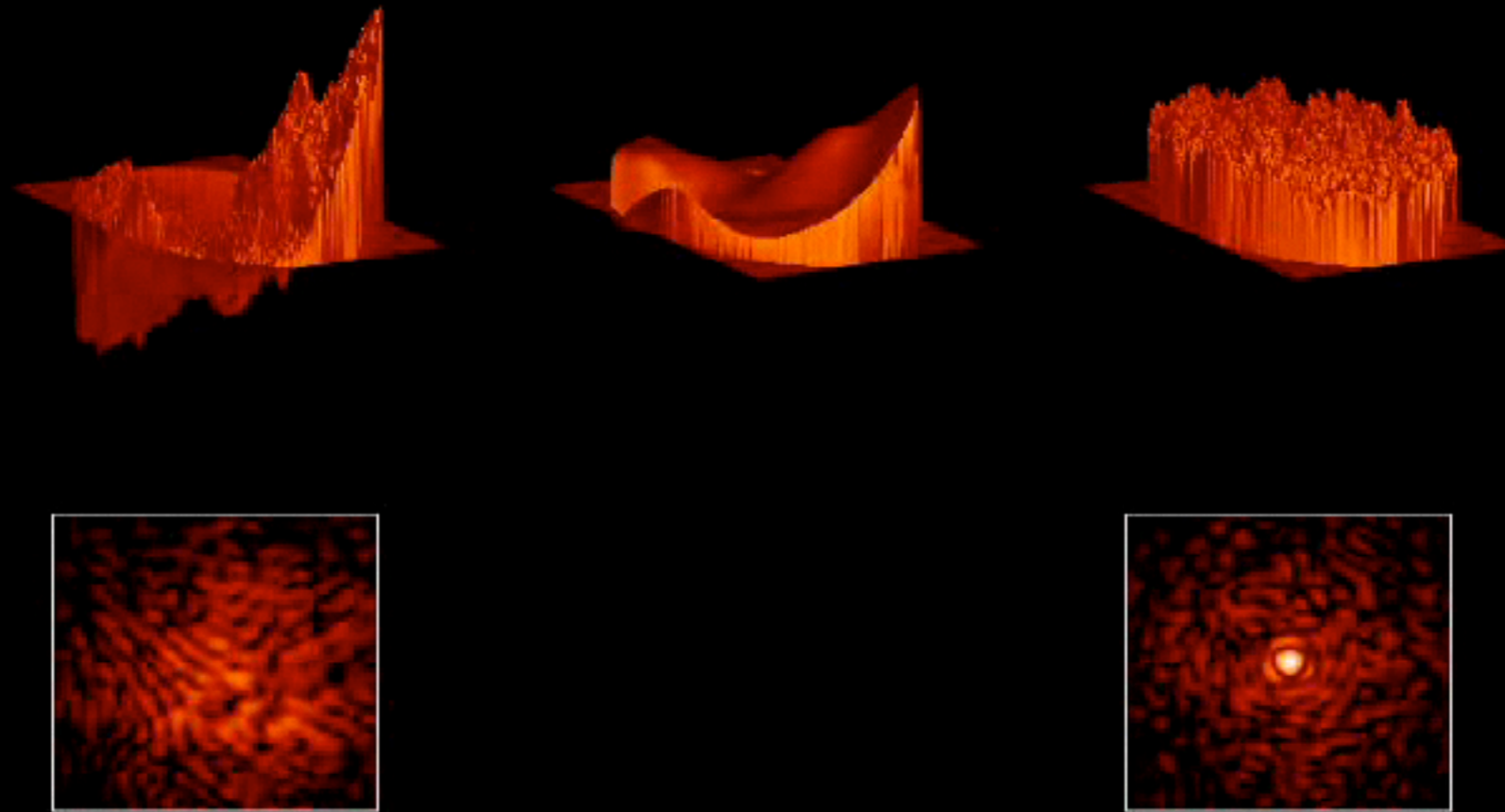




SCE_{Ex}AO: the Subaru Coronagraphic Extreme AO Project

Frantz Martinache
SCE_{Ex}AO Project Scientist
Subaru Telescope, NAOJ

Adaptive Optics: required but not sufficient!



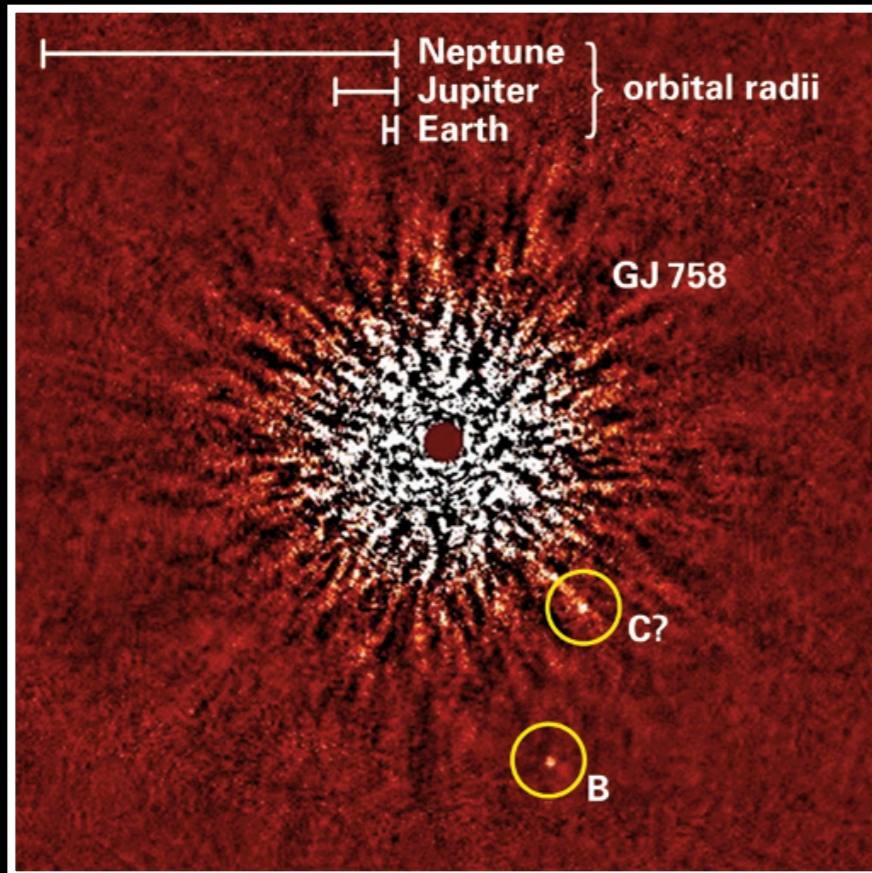
AO stabilizes the wavefront but **cannot** do a perfect job
You are at best left with *static or slowly varying aberrations* that
set contrast detection limits

Take care of an ill-posed problem

$$I = O \otimes \text{PSF}$$

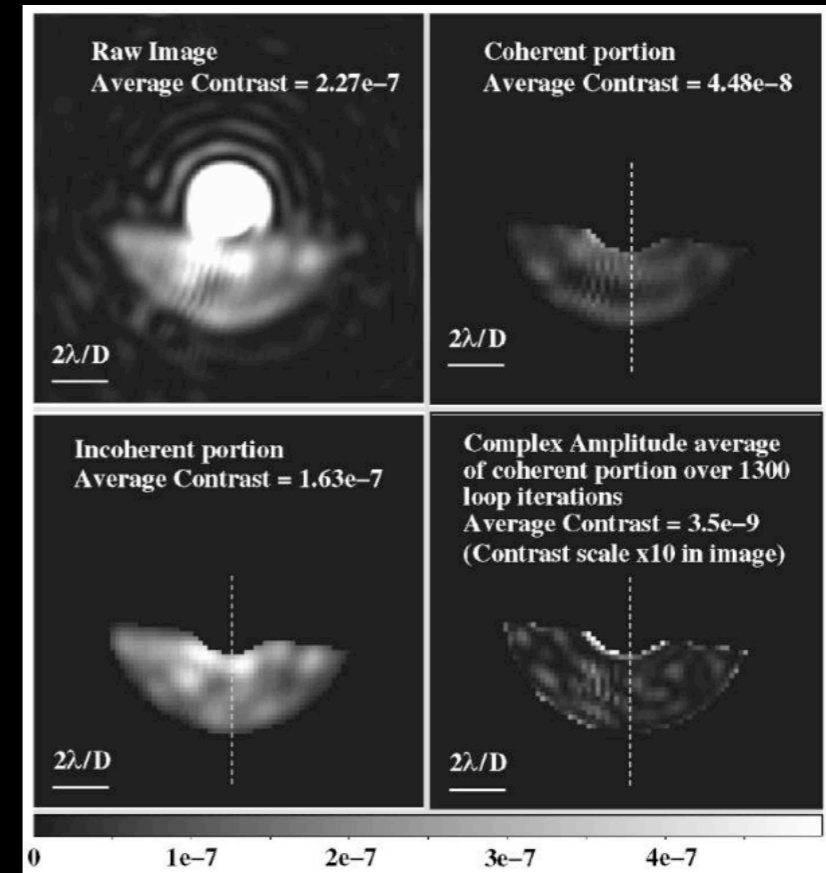
Eliminate the PSF out of the equation

the ADI way...



Marois et al, 2006, ApJ, 641, 556

the exAO way...

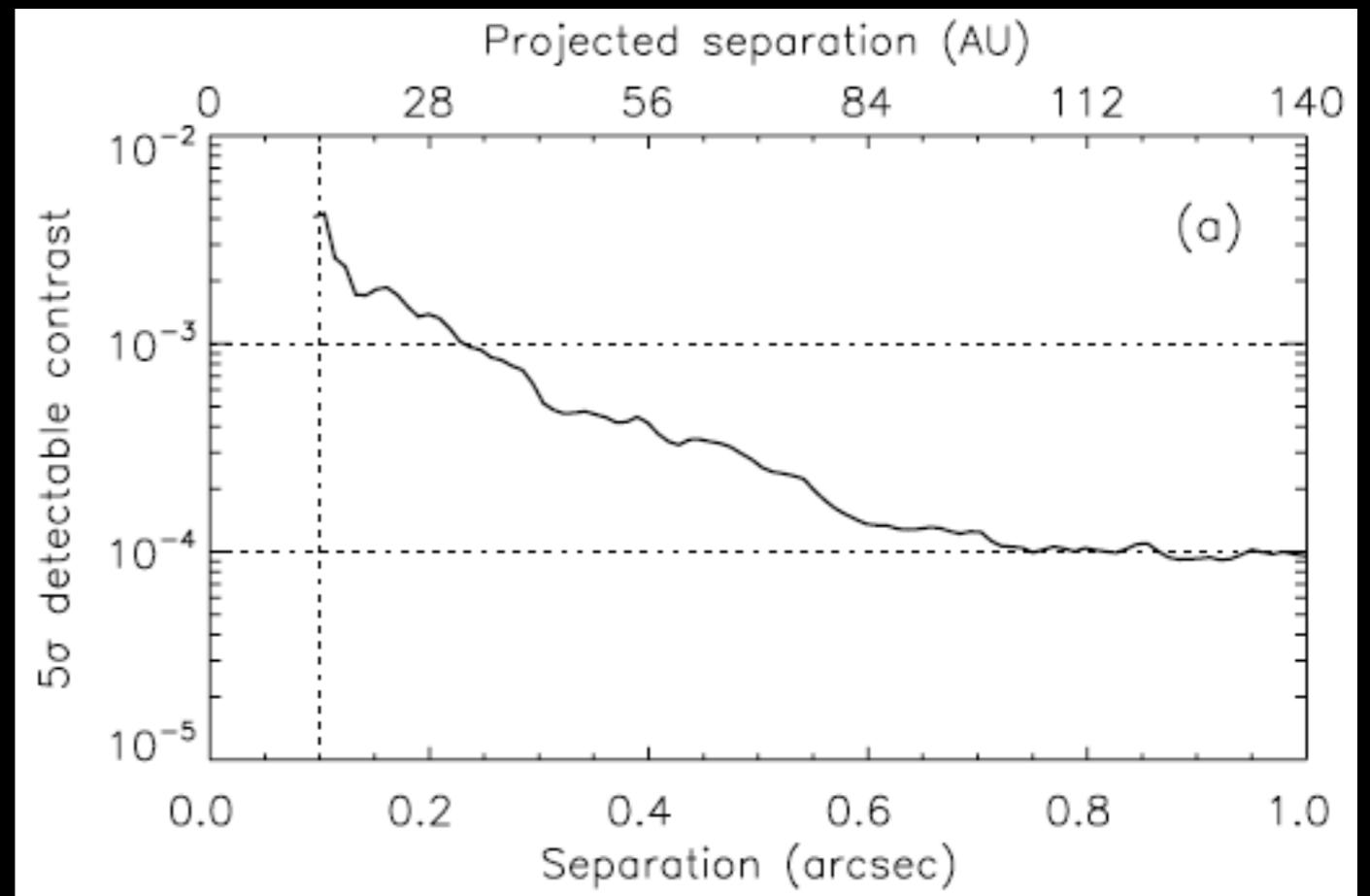


Guyon et al, 2009, PASP, 122, 71

SCEXAO

Angular Differential Imaging

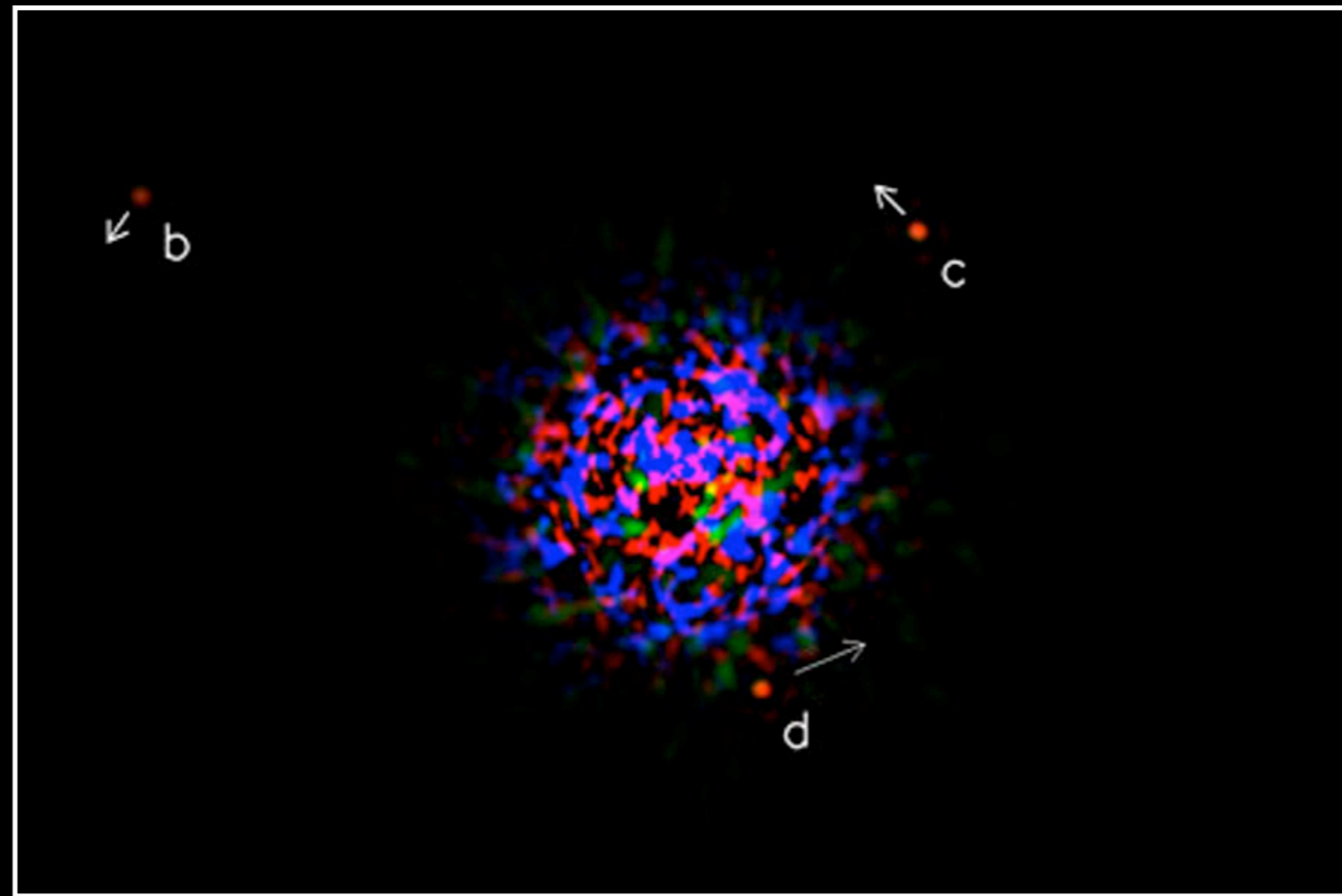
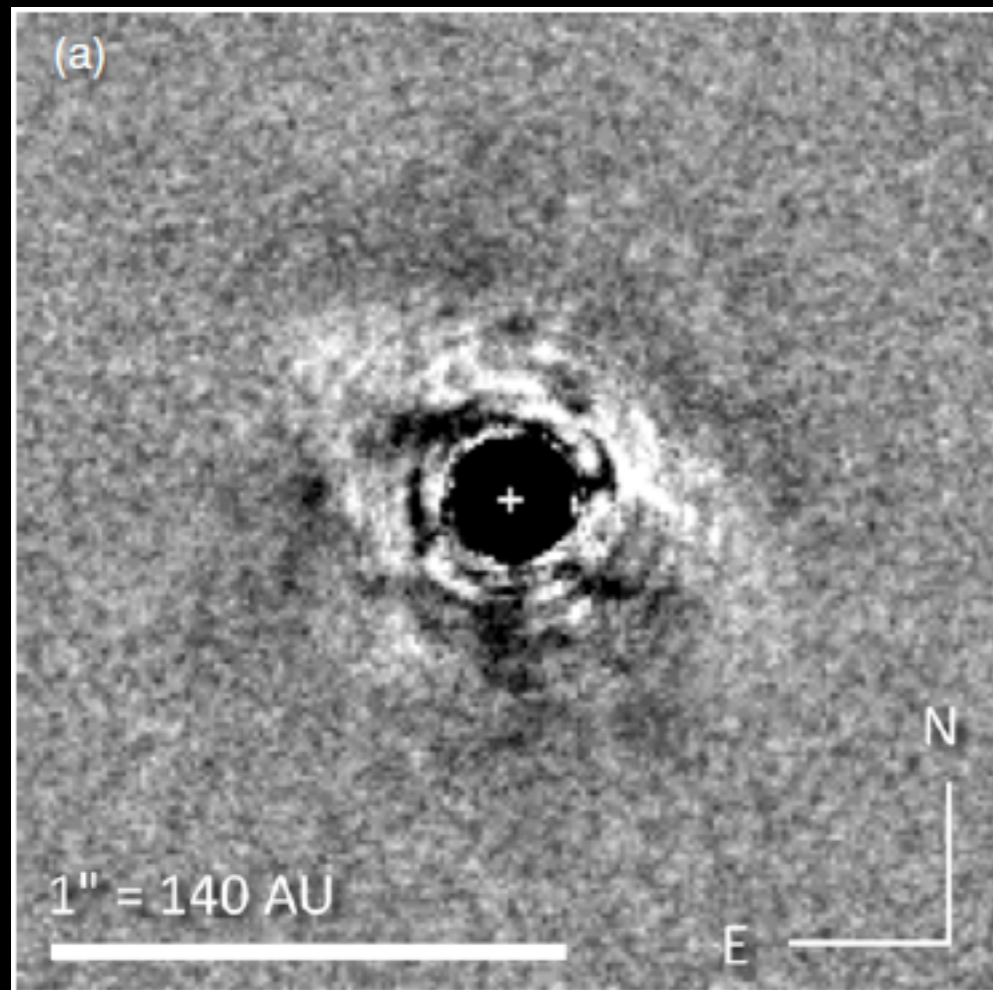
The “state of the art”:
optimized Angular
Differential Imaging
called LOCI



LOCI led to the obtention of direct images of sub-stellar companions around HR 8799, GJ 758, etc.

Remarkable... but reaches peak performance
somewhere around 0.5 arcsec

example of LOCI observations

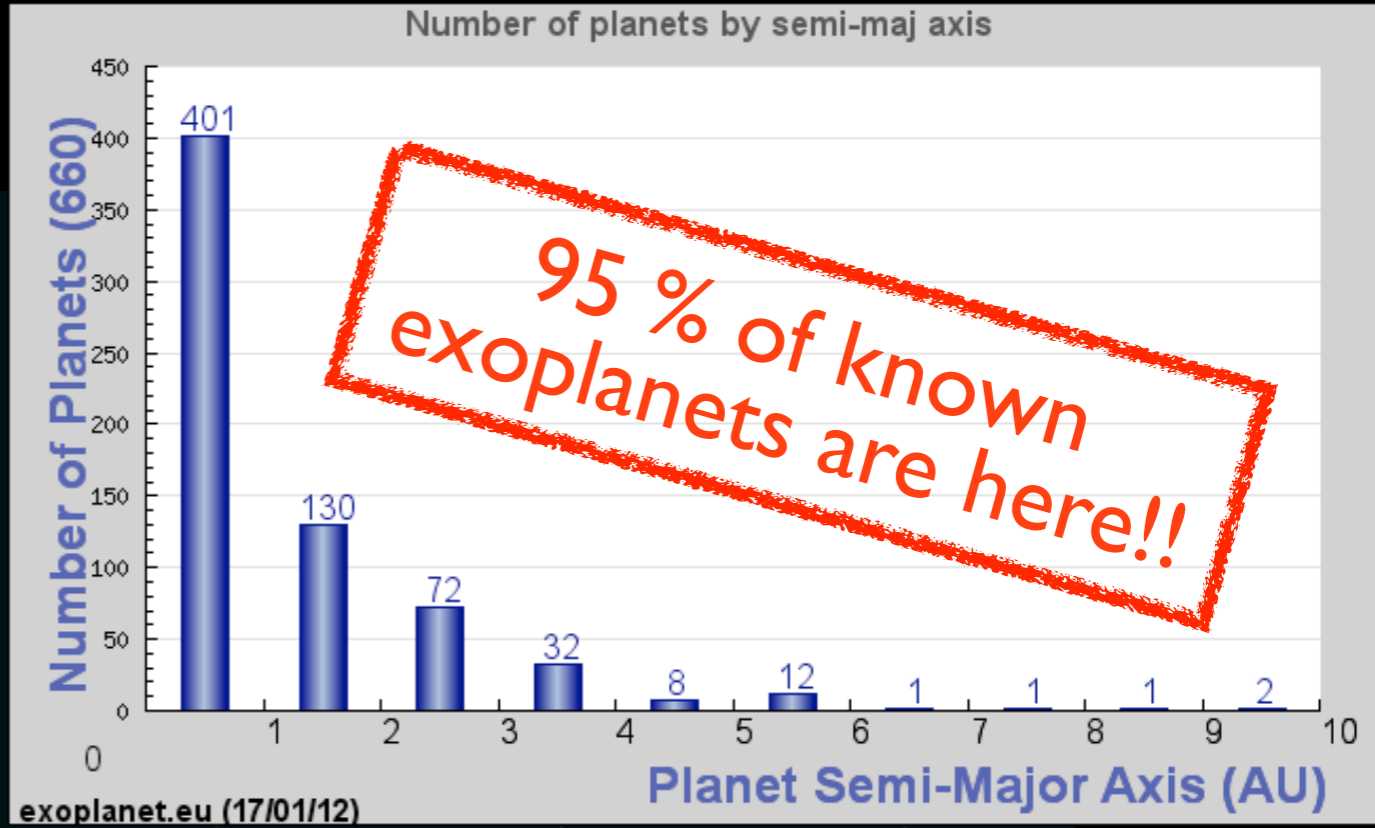
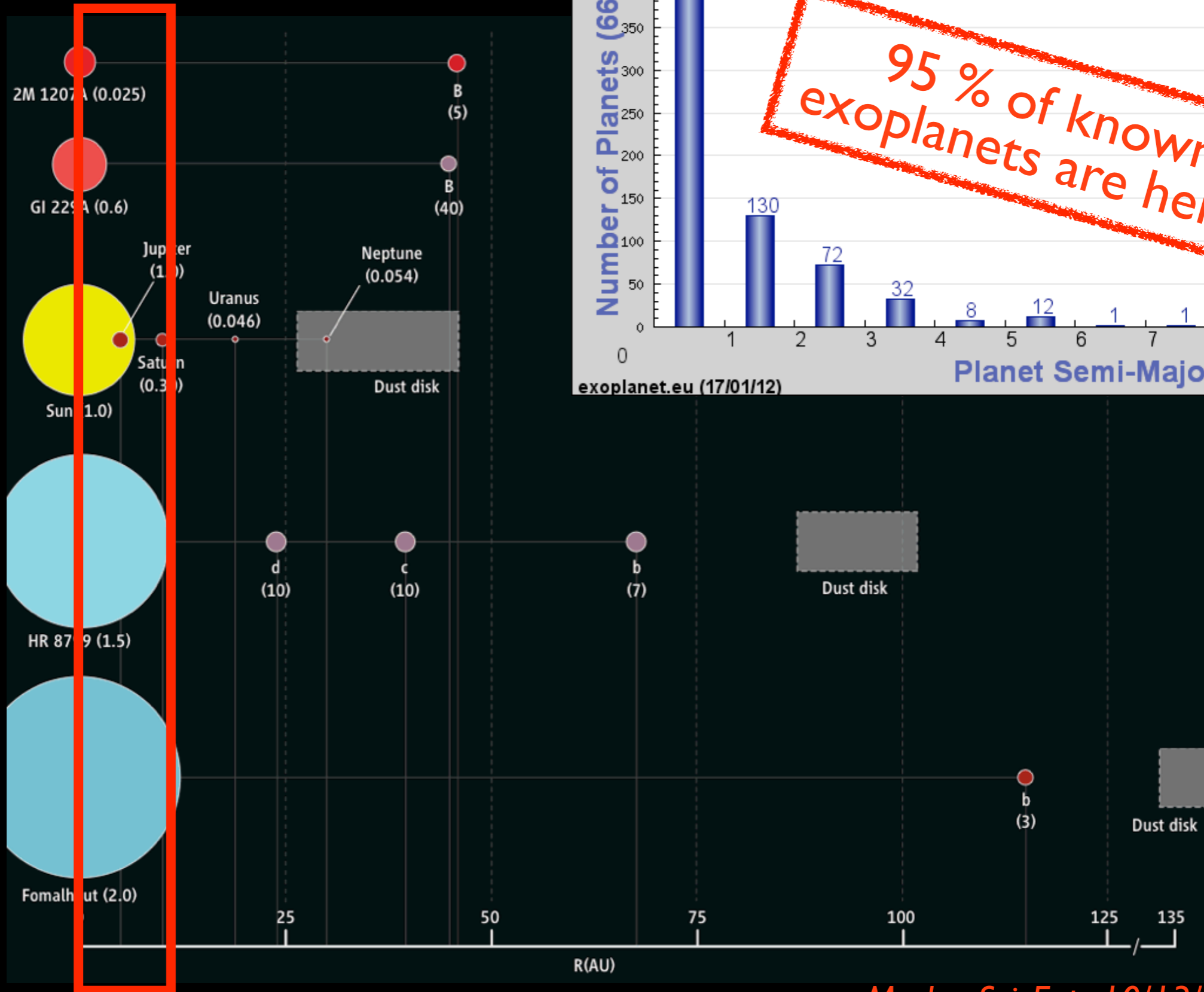


“conservative” LOCI image
of LkCa 15 by HiCIAO

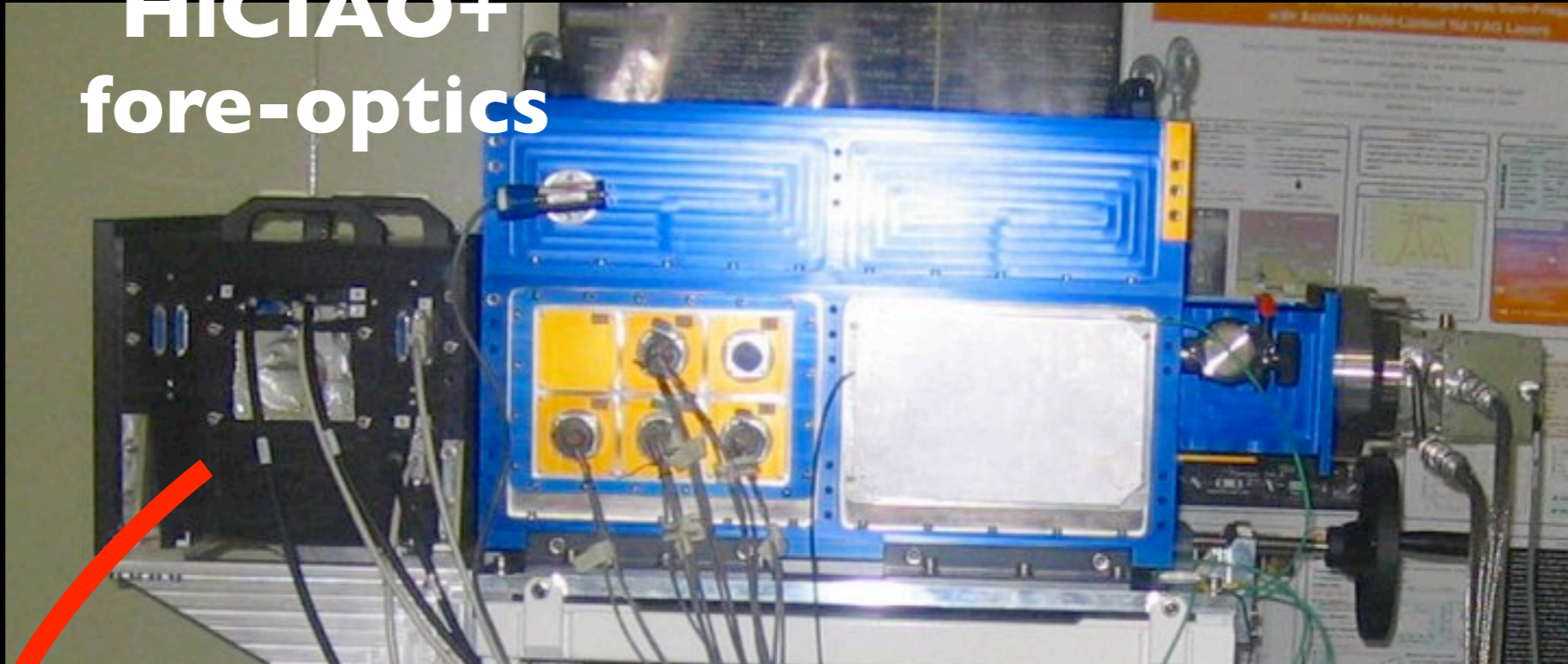
Thalman et al, 2010, ApJL, 718, 87

“iconic” LOCI image of the
planetary system orbiting HR 8799

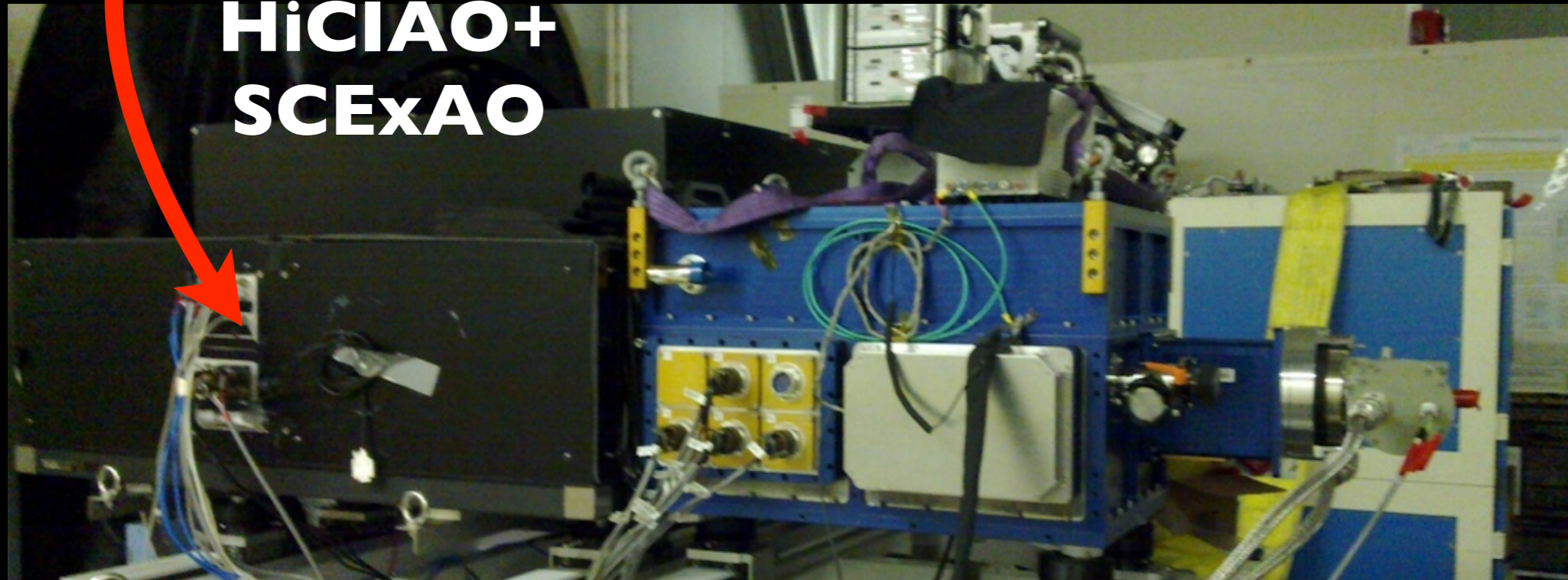
Marois et al, 2008, Science, 322, 1348

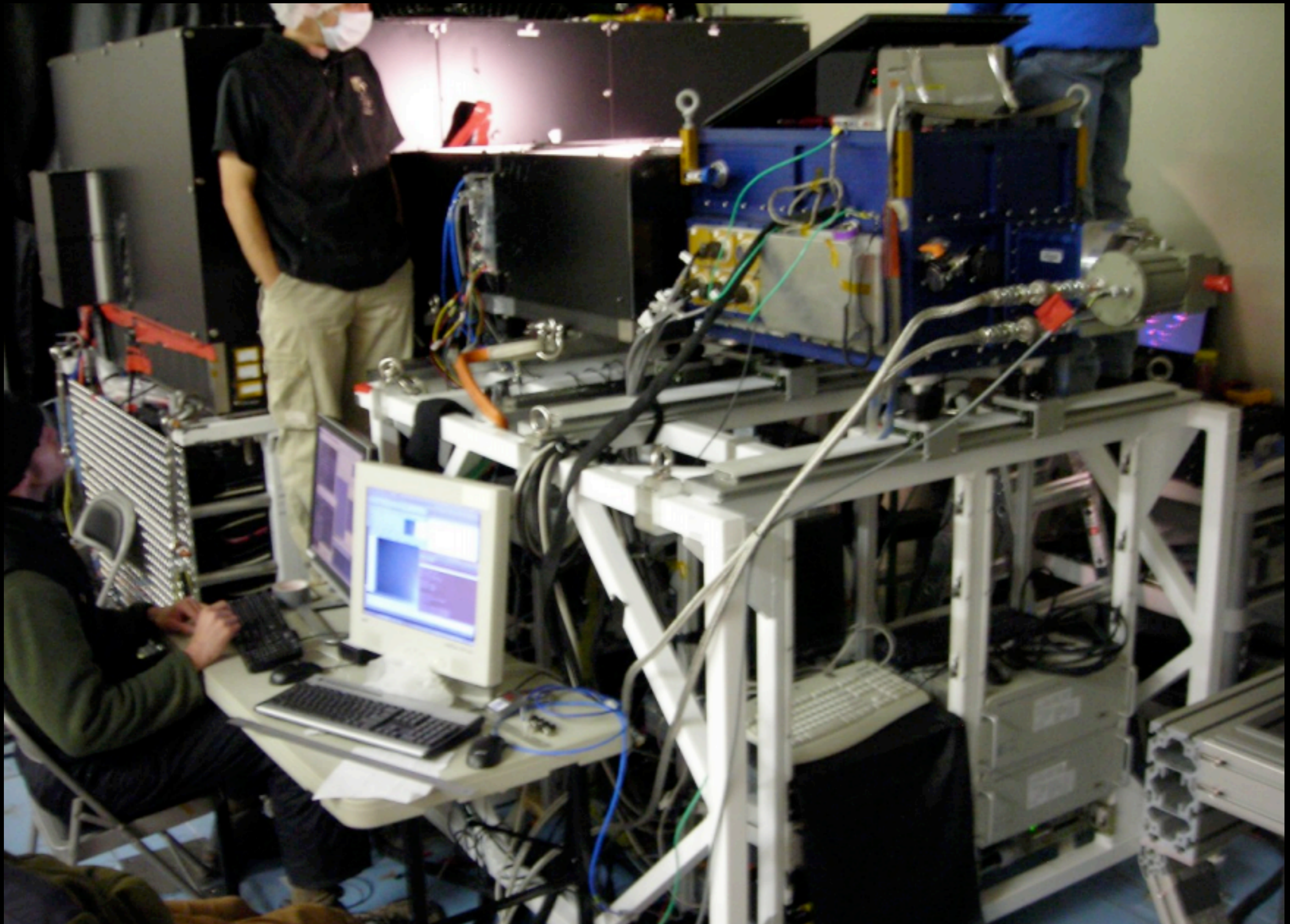


**HiCIAO+
fore-optics**



**HiCIAO+
SCEXAO**





SCExAO engineering run September 2011



**Internal
calib.
source**

**Visible imaging channel
soon to become
polarimetric NRM interferometry**

Pyramid WFS

**Common
Injection
module**

IR channel

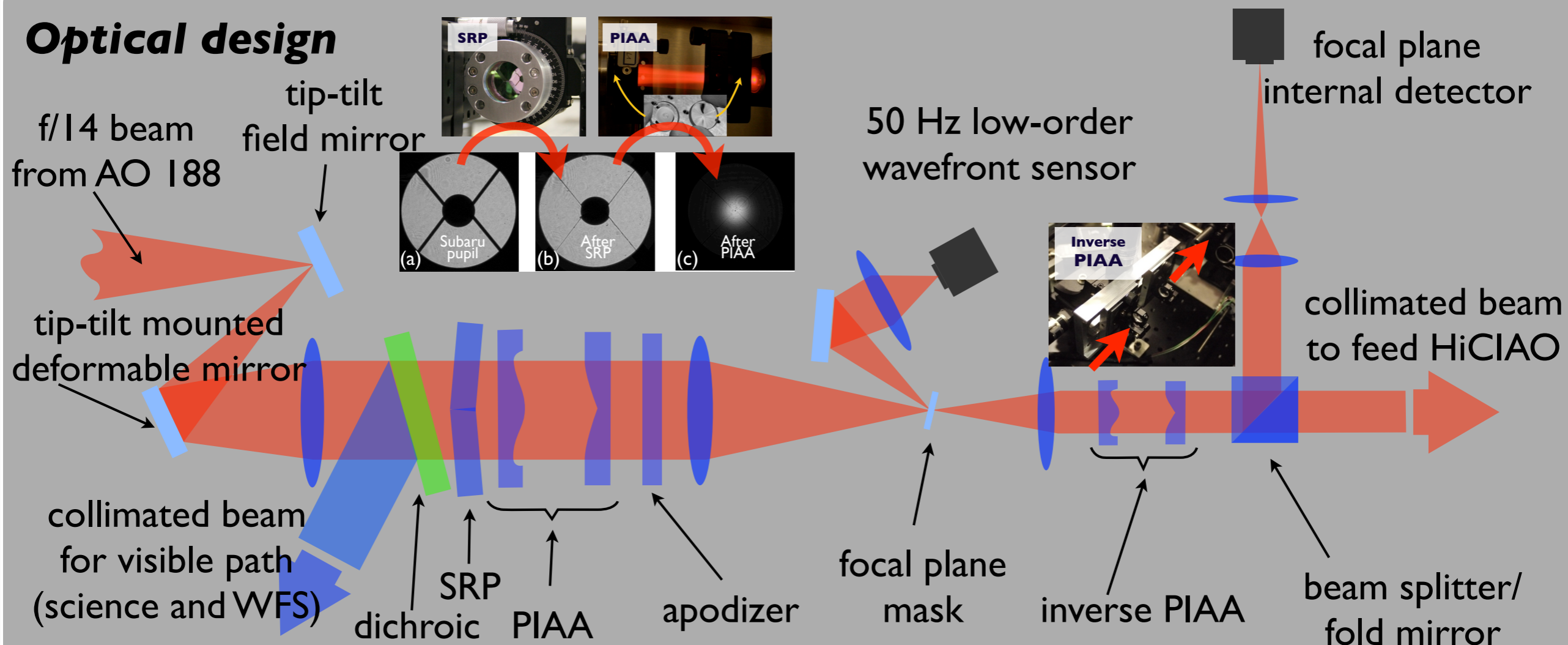
The SCExAO project

Taking advantage of a series of in house developments (PIAA, CLOWFS) and the existing infrastructure (AOI88 + HiCIAO) to put in the same box:

- a high-efficiency, high-performance **PIAA**-based coronagraph
- all the **calibration tools** we can think of now (wavefront sensors, active control of the focal plane image with a DM, post processing techniques, optional NRM)
- **simultaneous diffraction limited visible imager**

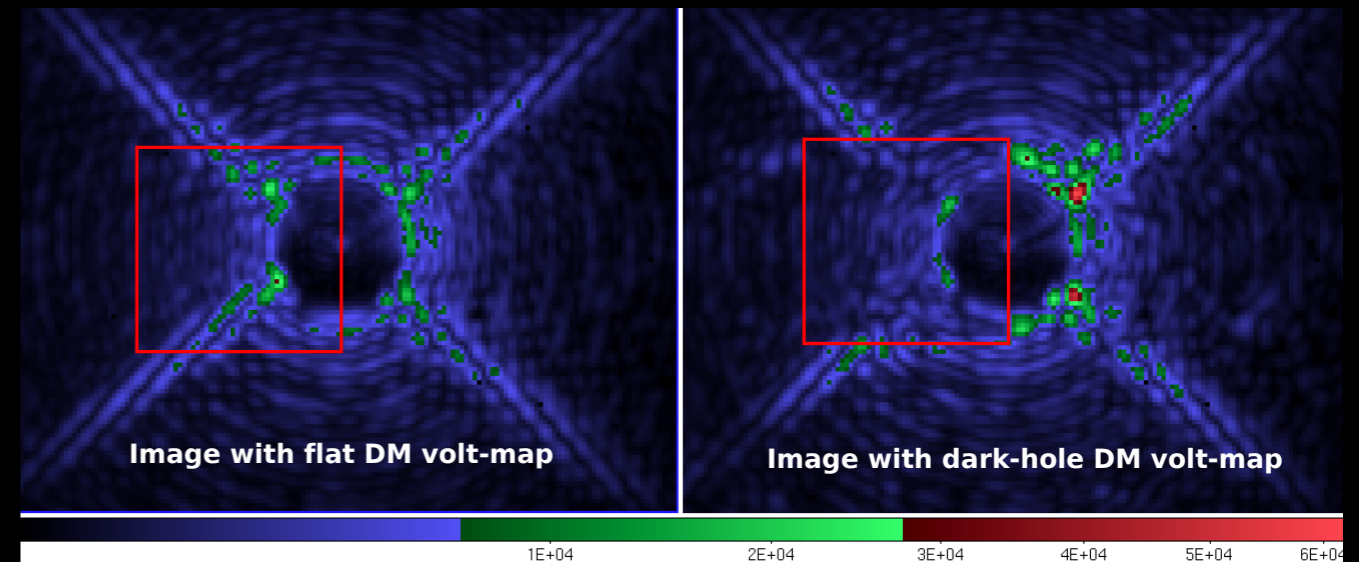
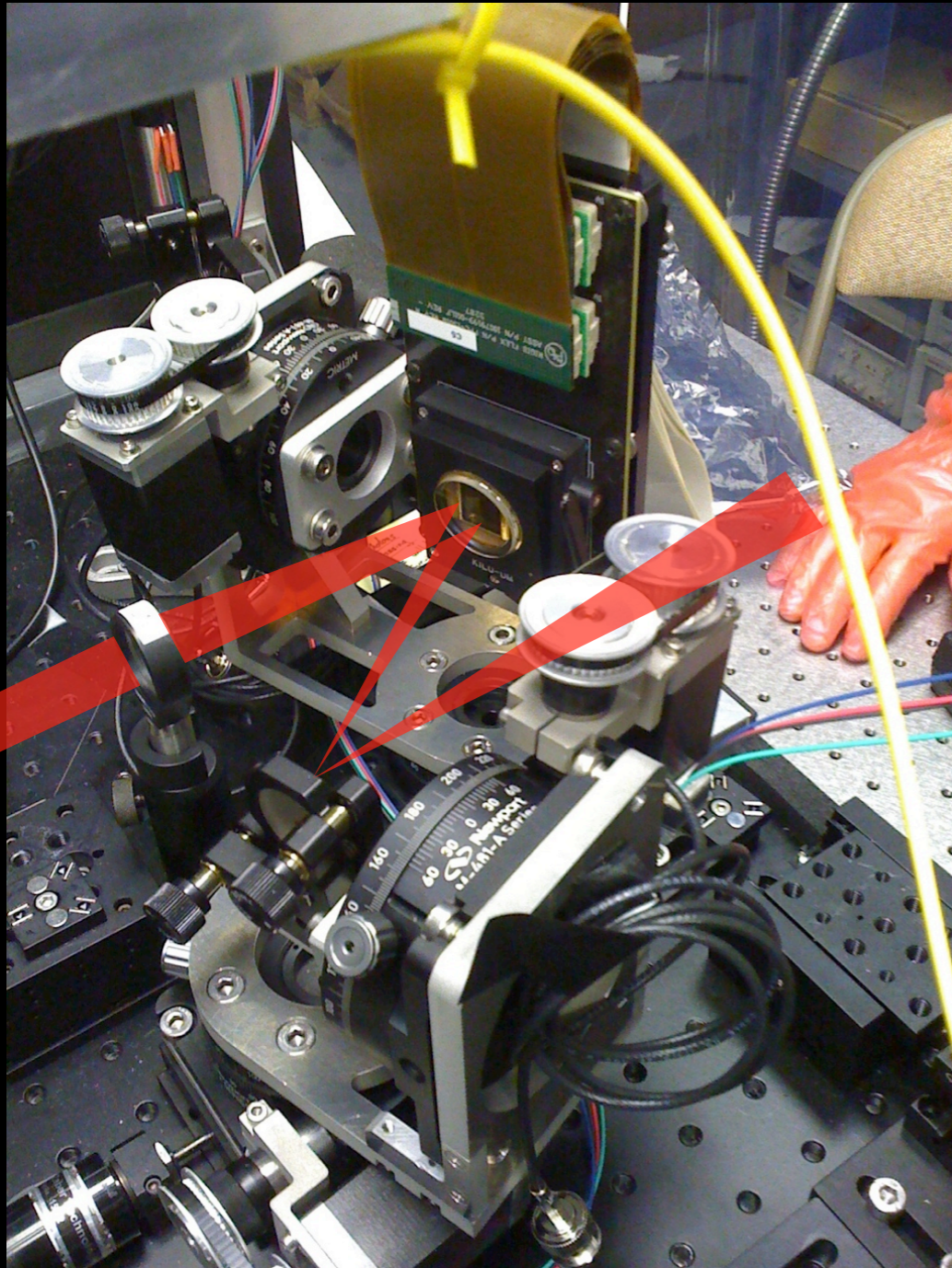
While keeping the design flexible for future improvements.

Optical design



Active speckle control

Instead of using a **passive approach**, and wait for the sky rotation like in ADI, use a DM with many actuators (~ 1000) gives **active control** of the speckles

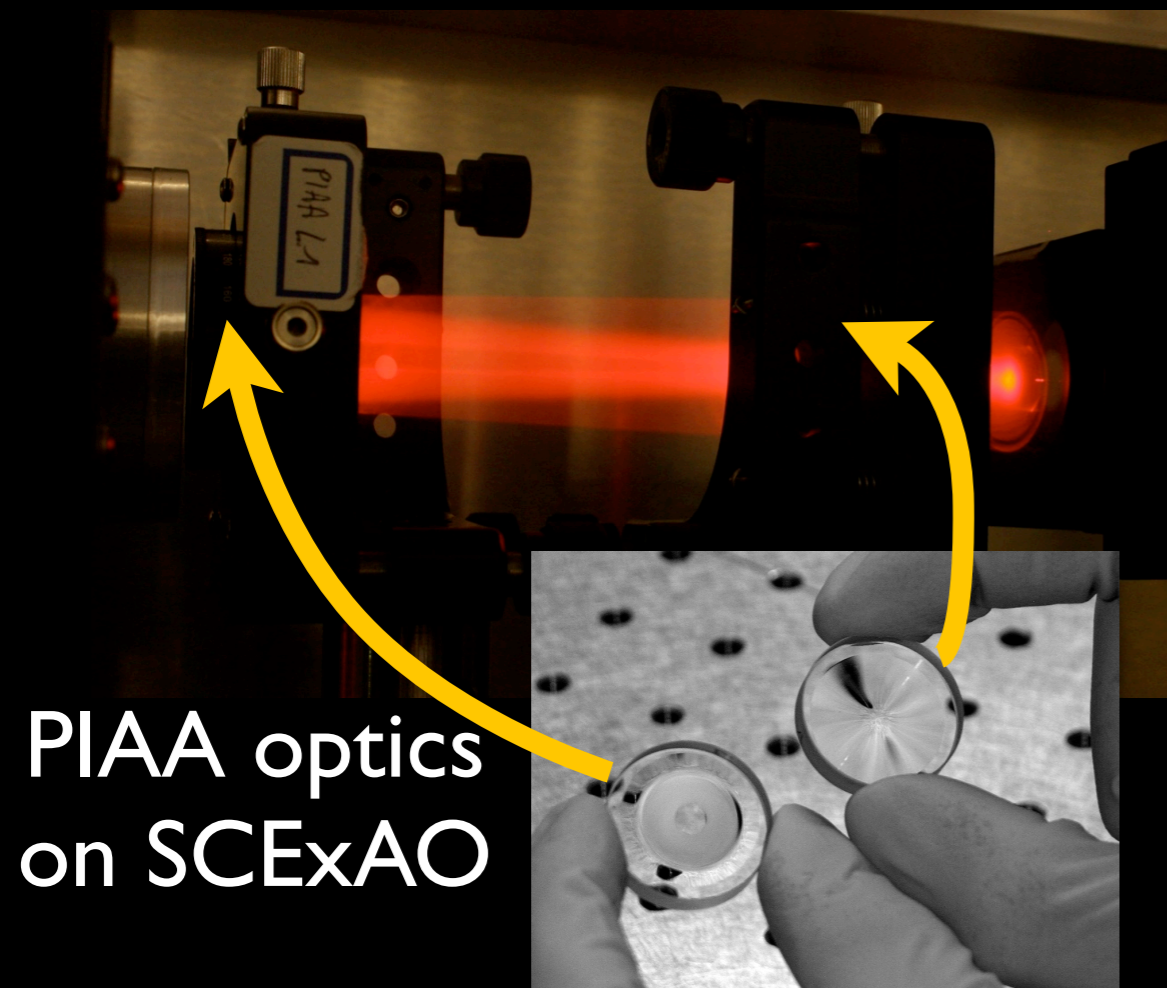


Strategy shared with comparable projects (GPI, SPHERE).

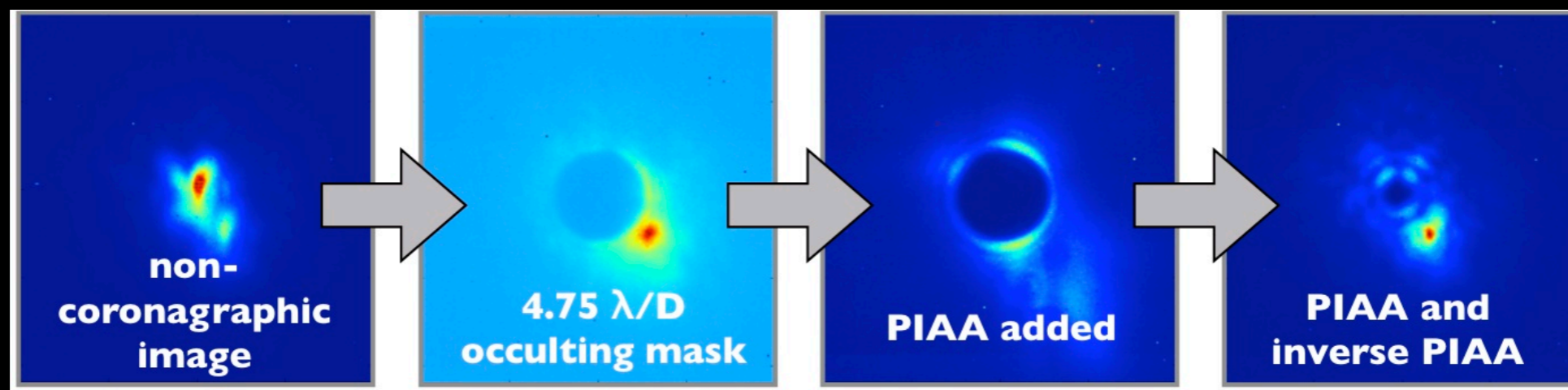
The true advantage of SCExAO is its **small IWA**

*Martinache & Guyon, 2011, AO4ELT
On sky results soon!*

High contrast @ 1λ with PIAA



Martinache & Guyon, 2011, AO4ELT



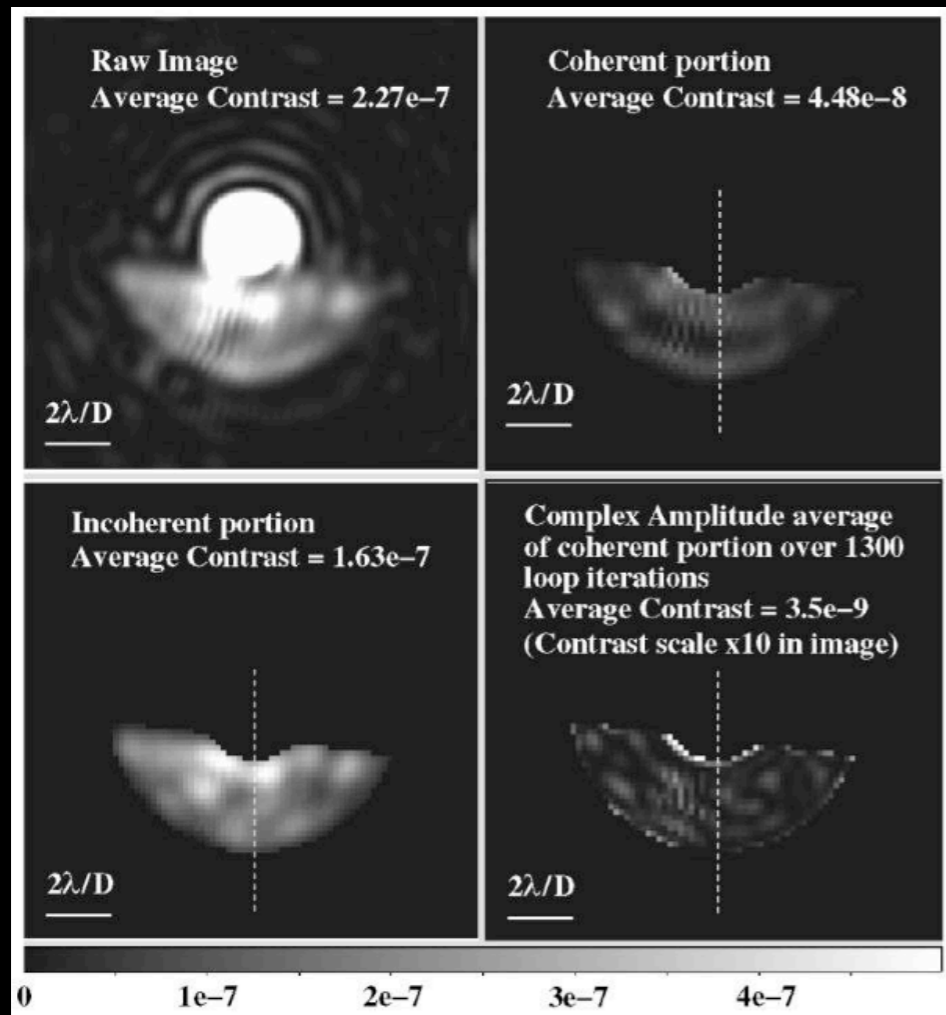
Images by HiCIAO

Active speckle control

Example of PIAA-based coronagraph high contrast results (laboratory):

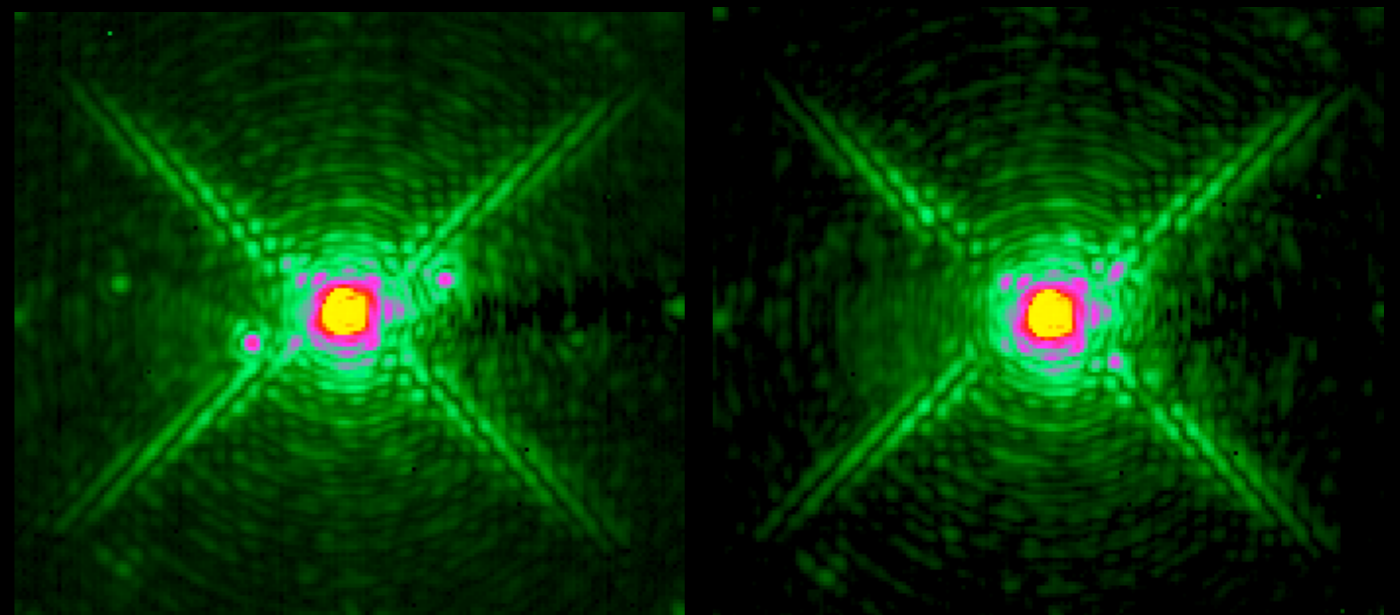
DM diversity identifies “coherent” fraction of the light in control FOV. Gains two orders of magnitude over raw contrast

Guyon et al, 2010, PASP, 122, 71



DM wavefront diversity in the presence of turbulence with SCExAO

Martinache et al, 2012, in prep



The background of the slide is a composite astronomical image. It features a bright star at the center, surrounded by a large, dark, and somewhat irregularly shaped protoplanetary disk. The disk has a reddish-brown hue, suggesting dust and gas. To the right of the star, a small, reddish-brown planet is visible, orbiting the star. The overall scene is set against a dark, star-filled background.

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