

GaiaNIR

ESA-led NIR all-sky astrometry mission

Daisuke Kawata

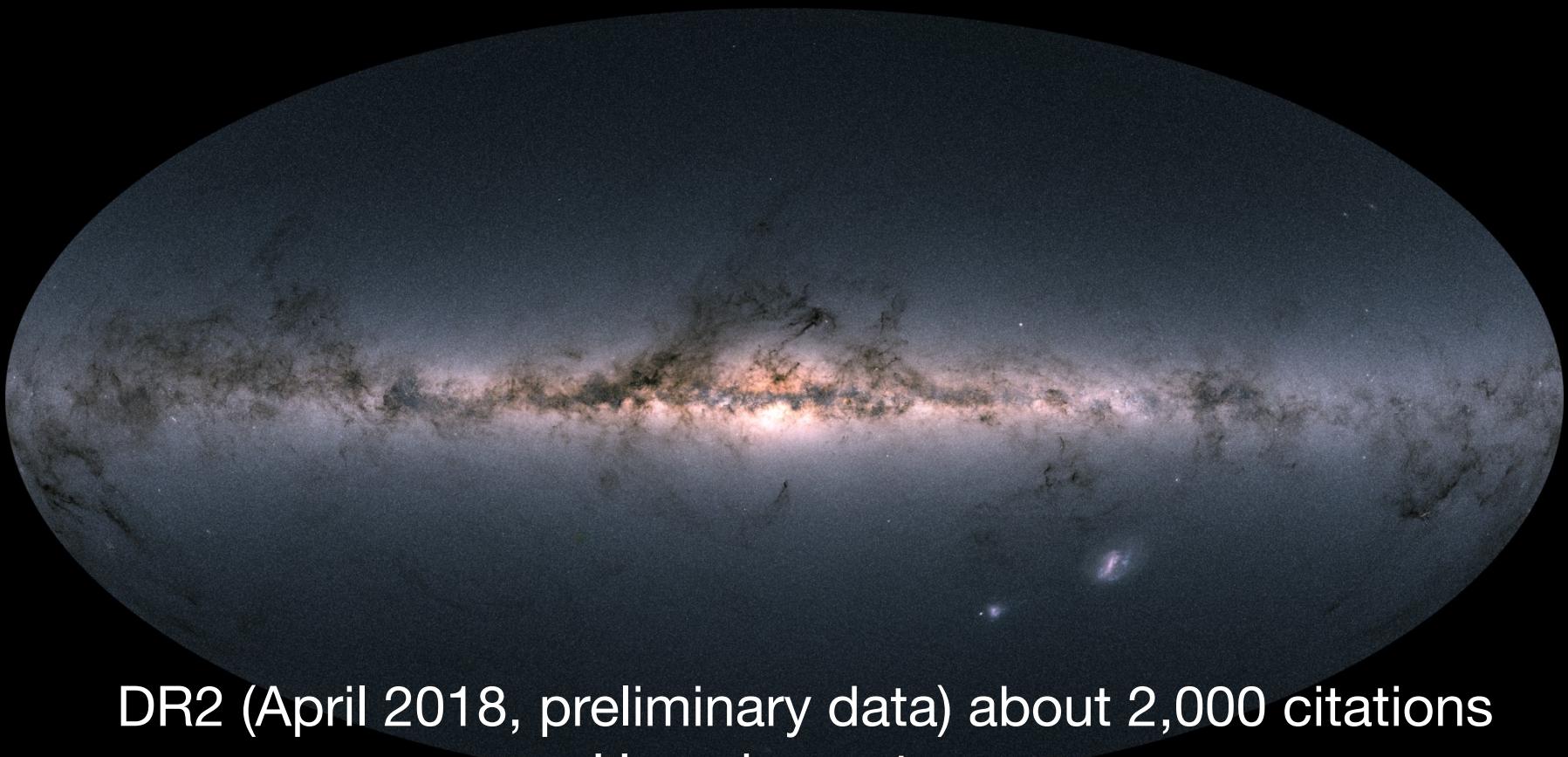
(Mullard Space Science Laboratory, University College London,
National Astronomical Observatory of Japan (NAOJ))

and 22 LoI co-Investigators

David Hobbs (Lund university, Sweden): PI of GaiaNIR

Junichi Baba (NAOJ), Masashi Chiba (Tohoku), Michiko Fujii (Tokyo),
Naoteru Gouda (NAOJ), Kohei Hattori (Michigan), Kohei Hayashi (ICRR,
Tokyo), Hirokazu Kataza (ISAS), Miho Ishigaki (Tohoku), Chiaki Kobayashi
(Hertfordshire, UK), **Noriyuki Matsunaga (Tokyo)**, Tadafumi Matsuno
(NAOJ), Makoto Miyoshi (NAOJ), Ryoichi Nishi (Niigata), Shogo Nishiyama
(Miyagi Education), Sakurako Okamoto (NAOJ), Alex Pettitt (Hokkaido),
Takahiro Sumi (Osaka), Masahiro Takada (IPMU, Tokyo), Takuji Tsujimoto
(NAOJ), Yoshiyuki Yamada (Kyoto), Taihei Yano (NAOJ)

ESA Gaia mission (2013-2022) Optical All-Sky Astrometry mission



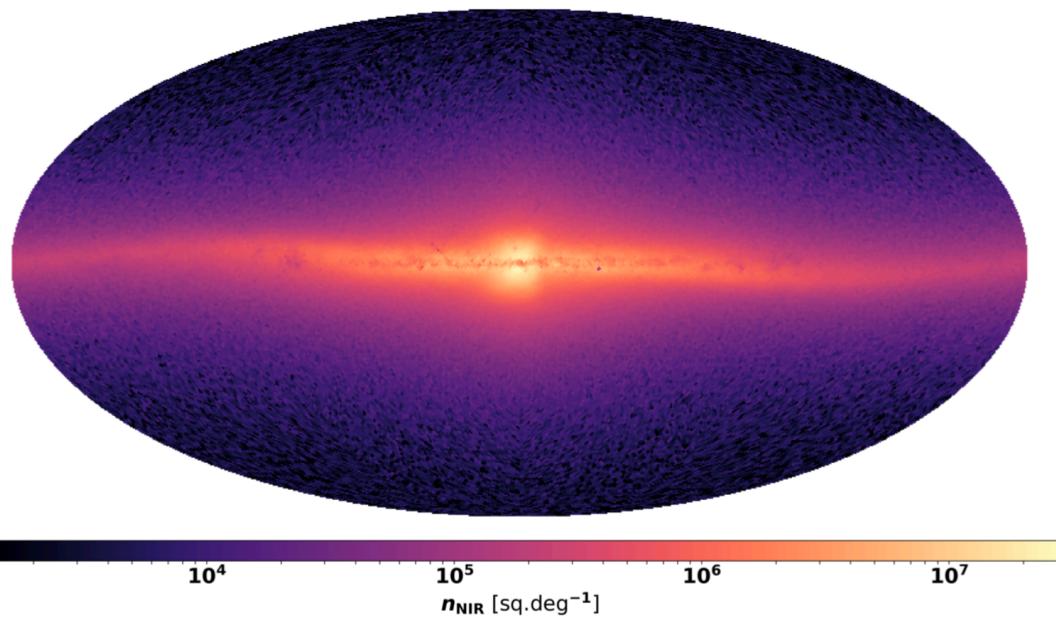
DR2 (April 2018, preliminary data) about 2,000 citations
Huge impacts on
from solar system, exoplanet to Galaxy and cosmology!

Astrometry: parallax (distance) and proper motion
providing new dimensions of information

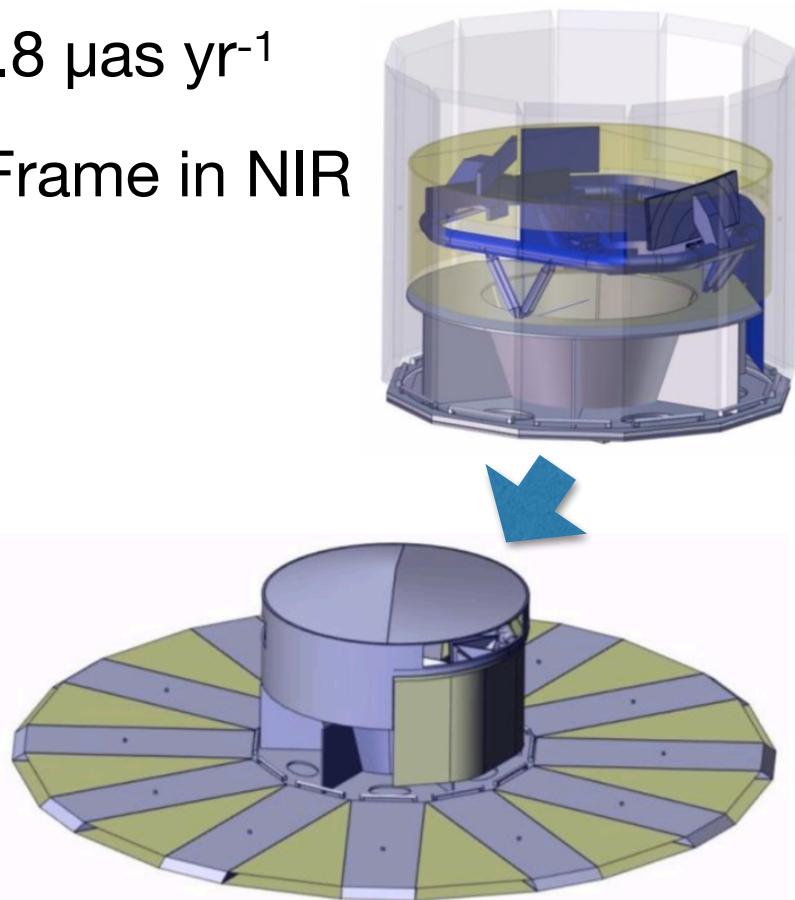
GaiaNIR

All-sky NIR astrometry in 2040s.

- NIR ($\lambda=0.8\text{-}1.8 \mu\text{m}$, $G_{\text{NIR}} < 20 \text{ mag}$)
Gaia-level astrometry ($\sim 10 \mu\text{as}$)
- 14 times better proper motion, $\sim 1.8 \mu\text{as yr}^{-1}$
- Update Gaia Celestial Reference Frame in NIR



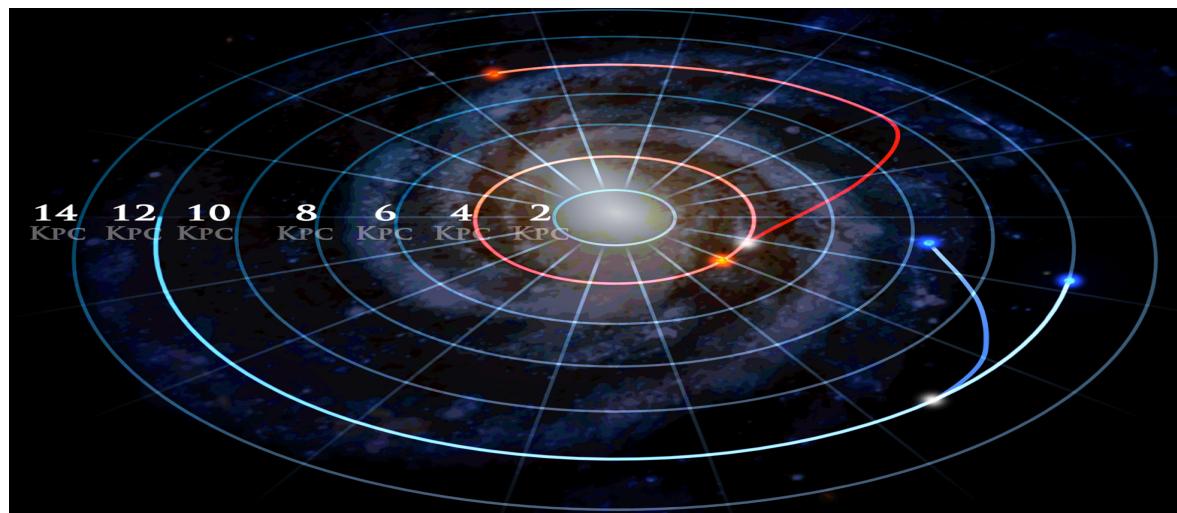
Gaia NIR sky (Hobbs+Kawata et al. 2019)



ESA CDF Study Report (2017)

NIR ($\lambda=0.8\text{-}1.8 \mu\text{m}$, $G_{\text{NIR}} < 20$ mag) Gaia-level astrometry ($\sim 10 \mu\text{as}$)
5 times more stars ($\sim 8\text{B}$ stars) than Gaia seeing through the dust

- Origin of the bar and spiral arms, wiggle in the disk
→ Radial Migration, Galactoseismology
Where are we from? The Sun's past orbit
Galactic Habitable Orbit?



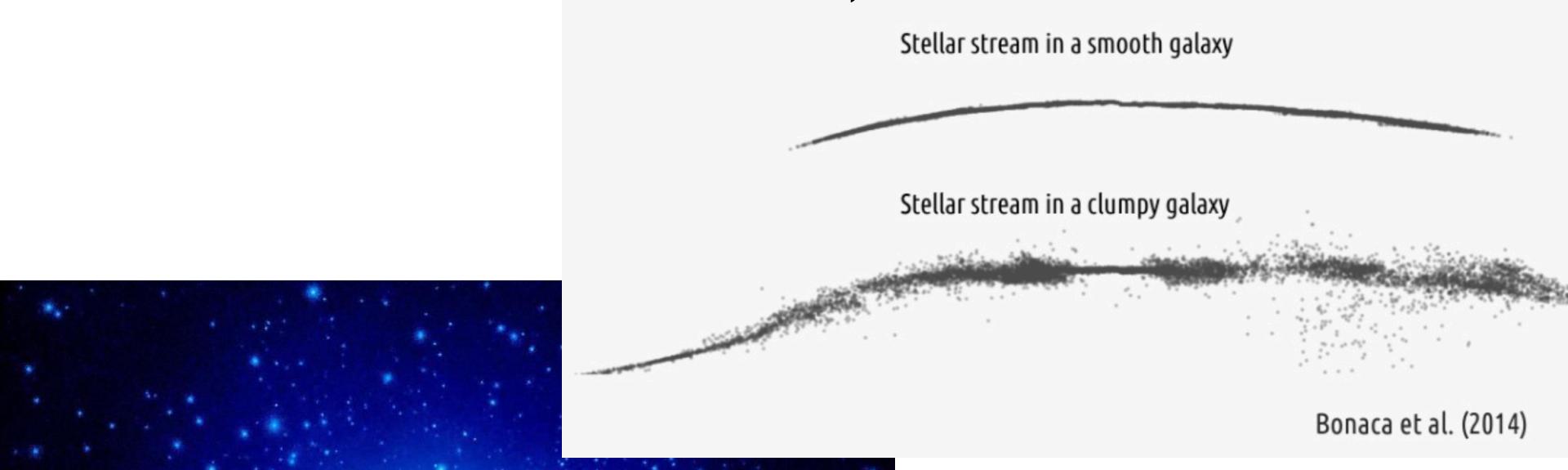
Credit:SDSS

- Plus more ... star forming regions
binaries (WD, BH and etc.) the distance to the GW sources
astrometric microlensing: primordial BHs as DM
Mira variable age, distance calibration
exoplanet of dusty young stars

~20 yrs time separation with Gaia and JASMINE

→ 14 times better proper motion $\sim 1.8 \mu\text{as yr}^{-1} \sim 10 \text{ km/s}$ at 1 Mpc
for ~1.6B stars common to Gaia, Galactic Centre stars to JASMINE

- Proper motion and internal stellar motion of Local Group galaxies, including LMC, SMC, M31, M33
- Gaps in stellar streams, wide binary populations
→ distribution of sub-DM halos, nature of DM



Improved proper motion

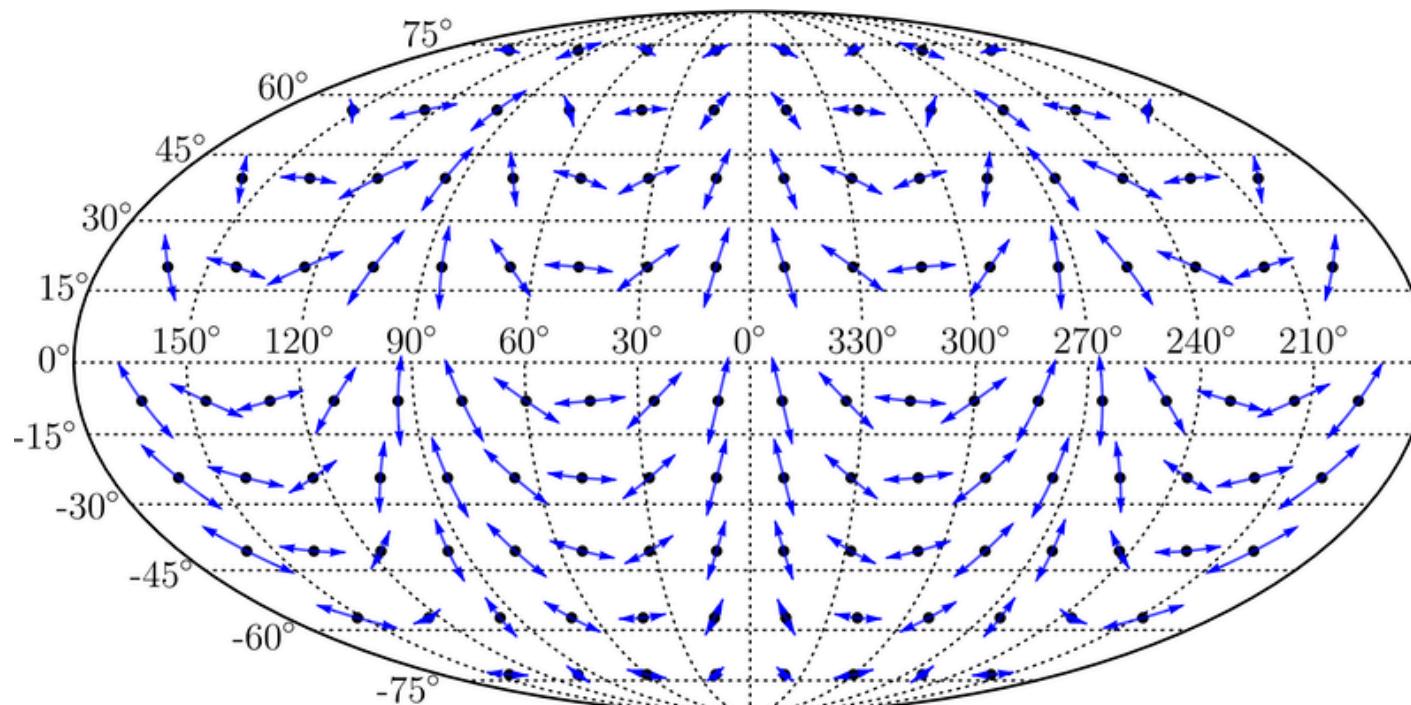
- Long period exo-planets
- Solar system:
acceleration of asteroids in the main asteroid belt,
understanding the origin of Near Earth Objects



Credit: KSJD

NIR Celestial Reference Frame in 2040s when Gaia CRF is degraded

- Reference frame for future (NIR) observation
- Astrometric gravitational wave
primordial gravitational wave?



Are you just waiting for the final catalogue? Why not contributing to the high impact mission!

- Japanese on-going legacy in Astrometry: e.g. VERA, JASMINE
(N)IR space astronomy: e.g. Subaru, PRIME, TAO, SPICA, TMT
Gravitational Wave: KAGRA, LiteBIRD
- Potential Japanese contribution
Launcher
Super-Super Invar (low expansion alloy, JASMINE)
Galactic centre JASMINE - GaiaNIR proper motion analysis
Subaru/PFS? spectroscopic follow up
- GaiaNIR status
Selected as one of 3 ESA's new science idea in 2017
ESA CDF completed in 2017 (284 page report)
Voyage 2050 WP submitted (w/Kawata, Gouda)
2040s launch?

